Sensitivity of future e<sup>+</sup>e<sup>-</sup> colliders to processes of dark matter production with light mediator exchange

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#### Snowmass 2021

EF10 DM@Colliders, topical meeting

October 27, 2021

A.F.Żarnecki (University of Warsaw) DM production with light mediator in e<sup>+</sup>e<sup>-</sup>

# **Motivation**

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# Probing Dark Matter with e<sup>+</sup>e<sup>-</sup>

# **Mono-photon signature**

The mono-photon signature is considered to be the most general way to look for DM particle production in future  $e^+e^-$  colliders.



DM can be pair produced in the  $e^+e^-$  collisions via exchange of a new mediator particle, which couples to both electrons (SM) and DM states

This process can be detected, if additional hard photon radiation from the initial state is observed in the detector...



# New analysis approach

Most of the studies performed so far focused on heavy mediator exchange (EFT limit) and coupling values  $\mathcal{O}(1)$  $\Rightarrow$  extracted were limits on DM or mediator masses

# In our study:

- focus on light mediator exchange (DM even lighter)
- consider very small mediator couplings to SM,  $\Gamma_{SM} \ll \Gamma_{tot}$



ILD study: arXiv:2001.03011 Phys. Rev. D 101, 075053 (2020)

CLIC study: arXiv:2103.06006

avial-vector

--- scalar

"Experimental-like" approach  $\Rightarrow$  focus on cross section limits as a function of mediator mass and width

A<sub>95</sub> [GeV]

2000

1000

H20 (500GeV



## Snowmass'2021 Lol

#### The proposal for this study was submitted in August 2020: SNOWMASS21-EF10\_EF9\_Filip\_Zarnecki-054.pdf

#### Final results have been just accepted for publication in EPJ C

arXiv:2107.11194

#### New approach to DM searches with mono-photon signature

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#### 1. MOTIVATION

One of the important goals of the proposed future e<sup>+</sup>e<sup>-</sup> collider experiments is the search for dark matter particles using different experimental approaches. The most general one is based on the mono-photon signature, which is expected when production of the invisible final state is accompanied by a hard photon from initial state radiation. We proposed the procedure<sup>1</sup> which allows for consistent, reliable simulation of mono-photon events in WHIZARD,<sup>2,3</sup> for both BSM signal and SM background processes, based on merging the matrix element calculations with the lepton ISR structure function.

#### 2. MERGING PROCEDURE

For precise kinematic description of photons entering the detector, we need to include hard photon emission directly in the process matrix element (ME) calculation. On the other hand, very soft and collinear photons should still be simulated with the parametric approach, taking into account proper summation of higher order corrections. A dedicated procedure for merging between the two regimes was proposed, exploiting variables<sup>1</sup>



# Outline



- 2 Simulating mono-photon events
- 3 Analysis framework





# Simulating mono-photon events

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# Simulating mono-photon events in $\operatorname{WHIZARD}$



For proper estimate of the mono-photon signature sensitivity consistent simulation of BSM processes and of the SM backgrounds is crucial.

"Irreducible" background comes from radiative neutrino pair-production



Detector acceptance & reconstruction efficiency

 $\Rightarrow$  significant contribution from radiative Bhabha scattering

WHIZARD provides the ISR structure function option that includes all orders of soft and soft-collinear photons as well as up to the third order in high-energy collinear photons.

However, WHIZARD ISR photons are not ordinary final state photons: they represent all photons radiated in the event from a given lepton line.

# Simulating mono-photon events in $\operatorname{WHIZARD}$



ISR structure function can not account for hard non-collinear photons  $\Rightarrow$  all "detectable" photons generated on Matrix Element level

Dedicated procedure developed to avoid double-counting of ISR and ME For details: J. Kalinowski et al., Eur. Phys. J. C 80 (2020) 634, arXiv:2004.14486

Two variables, calculated separately for each emitted photon:

$$\begin{array}{ll} q_{-} &=& \sqrt{4E_{0}E_{\gamma}}\cdot\sin\frac{\theta_{\gamma}}{2} \ , \\ q_{+} &=& \sqrt{4E_{0}E_{\gamma}}\cdot\cos\frac{\theta_{\gamma}}{2} \end{array}$$

are used to separate "soft ISR" emission region from the region described by ME calculations.





## Validation of the procedure

WHIZARD predictions were compared to the results from the KKMC code for  $e^+e^- \rightarrow \nu\bar{\nu} + N\gamma$  3 TeV CLIC



 $\Rightarrow$  very good agreement observed (both for shape and normalisation)

#### For more details:

J. Kalinowski et al., Eur. Phys. J. C 80 (2020) 634, arXiv:2004.14486



# Simplified DM model

UFO model covering most popular scenarios of DM pair-production Possible mediators:  $\Rightarrow$  Feynrules

- scalar
- pseudo-scalar
- vector
- pseudo-vector
- V-A coupling
- V+A coupling
- Possible DM candidates:
  - real or complex scalar
  - Majorana or Dirac fermion
  - real vector

Cross section for  $e^+e^- \rightarrow \chi\chi$  for  $M_{\chi} = 50 \text{ GeV}$  and  $M_{Y} = 300 \text{ GeV}$ 





# ISR rejection probability

Fraction of events generated by WHIZARD **removed** in merging procedure (ISR photons emitted in the phase-space region covered by ME)



CLIC @ 3 TeV





# **Tagging efficiency**

based on  $\operatorname{DELPHES}$  simulation

Mono-photons reconstructed only in a fraction of generated signal event

 $\sigma\left(e^{+}e^{-} \rightarrow \chi \; \chi \; \gamma_{_{\mathrm{tag}}}\right) \; = \; f_{\mathrm{mono-photon}} \cdot \sigma\left(e^{+}e^{-} \rightarrow \chi \; \chi \; (\gamma) \;\right)$ 

ILC @ 500 GeV

CLIC @ 3 TeV



# **Analysis framework**

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# Analysis framework



### **Event selection**

On generator level:

- 1, 2 or 3 ME photons nonradiative events for signal only (for normalisation)
- all ME photons with  $q_{\pm} > 1 \text{ GeV}$  &  $E^{\gamma} > 1 \text{ GeV}$ rejected are events with  $q_{\pm} > 1 \text{ GeV}$  &  $E^{\gamma} > 1 \text{ GeV}$ for any of the ISR photons
- at least one ME photon with  $p_T^{\gamma} > 2 \text{ GeV } \& 5^{\circ} < \theta^{\gamma} < 175^{\circ}$  (ILC 500 GeV)  $p_T^{\gamma} > 5 \text{ GeV } \& 7^{\circ} < \theta^{\gamma} < 173^{\circ}$  (CLIC 3 TeV)

Delphes framework used for detector simulation and event reconstruction.

### Require:

- single photon with  $p_T^{\gamma} > 3 \ GeV \& |\eta^{\gamma}| < 2.8 \ (ILC)$   $p_T^{\gamma} > 10 \ GeV \& |\eta^{\gamma}| < 2.6 \ (CLIC)$ 
  - no other activity in the detector other reconstructed objects
    - no electrons
    - no LumiCal photons
    - no BeamCal photons
    - no jets



## **Background distributions**

Two SM backgrounds considered: with up to 3 ME photons Bhabha scattering and (radiative) neutrino pair production





# **Background vs Signal distributions**



For mono-photon events, two variables fully describe event kinematics  $\Rightarrow$  use 2D distribution of  $(p_T^{\gamma}, \eta)$  to constrain DM production Background Signal





# Systematic uncertainties

following ILD study: Phys. Rev. D 101, 075053 (2020), arXiv:2001.03011

Considered sources of uncertainties:

- Integrated luminosity uncertainty of 0.26% uncorrelated between polarisations
- Luminosity spectra shape uncertainty correlated between polarisations
- Uncertainty in neutrino background normalisation of 0.2% (th+exp) correlated between polarisations
- Uncertainty in Bhabha background normalisation of 1% (th+exp) correlated between polarisations
- Uncertainty on beam polarisation of 0.02–0.08% (ILC)/0.2% (CLIC) correlated for runs with same beam polarisation at ILC

 $\Rightarrow$  nuisance parameters in the model fit (11 for ILC, 7 for CLIC)

# Results

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Cross section limits

for radiative events (with tagged photon)

Vector Mediator

 $\Gamma/M = 0.03$  wi

with and without systematics

### ILC @ 500 GeV

CLIC @ 3 TeV



Systematic effects reduced for on-shell production of narrow mediator



Cross section limits

for radiative events (with tagged photon)

Vector Mediator

 $\Gamma/M = 0.5$ 

with and without systematics

### ILC @ 500 GeV

CLIC @ 3 TeV



Systematic effects reduced for on-shell production of narrow mediator



Cross section limitsfor radiative events (with tagged photon)Impact of beam polarisationassuming  $4 \text{ ab}^{-1}$  for ILC @ 500 GeV

#### Vector mediator

Scalar mediator



Combining four data sets taken with different beam polarisation settings significantly reduces impact of systematic uncertainties...

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## Impact of systematic uncertainties

How important are the external constraints on the systematic effects? Eg. precision of the lumionosity measurement or theoretical calculations...

Total cross section limits for Vector mediator with  $M_Y=2$  TeV,  $\Gamma/M=0.03$ 

Single uncertainty varied

All systematics (one varied)



Most of the systemactic effects are constrained by the data itself!!!



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**Cross section limits** for total DM production cross section Corrected for probability of hard photon tagging!

Combined limits for Vector mediator

ILC @ 500 GeV

CLIC @ 3 TeV



Radiation suppressed for narrow mediator with  $M_Y \sim \sqrt{s} \Rightarrow$  weaker limits



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**Cross section limits** for total DM production cross section Corrected for probability of hard photon tagging!

Combined limits for mediators with  $\Gamma/M = 0.03$ 

II C @ 500 GeV

CLIC @ 3 TeV



Radiation suppressed for narrow mediator with  $M_Y \sim \sqrt{s} \Rightarrow$  weaker limits

# **Coupling limits**

#### Combined limits for Vector mediator



Almost uniform sensitivity to mediator coupling  $g_{eeY}$  up to kinematic limit.

ILC @ 500 GeV

# **Coupling limits**

### Combined limits for mediators with $\Gamma/M = 0.03$

#### g<sub>ee7</sub> (95%CL) (95%CL ILC 10 CLIC 4 ab<sup>-1</sup> @ 500 GeV 5 ab<sup>-1</sup> @ 3 TeV /-A coupling V-A coupling $g_{\rm ee\gamma}$ V+A coupling V+A coupling 10<sup>-1</sup> 10<sup>-2</sup> 10<sup>-2</sup> 10<sup>-3</sup> 10<sup>3</sup> 10<sup>2</sup> 10<sup>2</sup> 10<sup>3</sup> 10<sup>4</sup> M<sub>v</sub> [GeV] M<sub>v</sub> [GeV]

CLIC @ 3 TeV

Almost uniform sensitivity to mediator coupling  $g_{eeY}$  up to kinematic limit.







Effective mass scale limits  $\Lambda^2 = \frac{M_Y^2}{|g_{eeY}g_{yyY}|}$ 

Combined limits for Vector mediator

ILC @ 500 GeV

CLIC @ 3 TeV





Effective mass scale limits

 $\Lambda^2 = \frac{\mathsf{M}_Y^2}{|\mathsf{g}_{eeY}\mathsf{g}_{\chi\chi Y}|}$ 

Combined limits for mediators with  $\Gamma/M=0.03$ 

ILC @ 500 GeV

CLIC @ 3 TeV



# Conclusions

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# Sensitivity of future linear $e^+e^-$ colliders to processes of dark matter production with light mediator exchange

arXiv:2107.11194, in print

Mono-photon signature: the most general way to look for DM production, EFT sensitivity extending to the O(10) TeV mass scales

New framework for mono-photon analysis developed focus on light mediator exchange and very small mediator couplings to SM

•  $\mathcal{O}(1\,{
m fb})$  limits on the radiative production  $e^+e^- o \chi\chi\gamma_{
m tag}$ 

•  $\mathcal{O}(10 \text{ fb})$  limits on the DM pair-production  $e^+e^- \rightarrow \chi \chi(\gamma)$ except for the resonance region  $M_Y \sim \sqrt{s}$ 

•  $\mathcal{O}(10^{-3}-10^{-2})$  limits on the mediator coupling to electrons up to the kinematic limit  $M_Y \leq \sqrt{s}$ 

For light mediators limits more stringent than those expected from direct resonance search in SM decay channels

# Thank you!

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## **Running scenarios**

Staged construction assumed for both ILC and CLIC. Results presented in this talk focus on the highest energy stages.

## ILC

Total of  $4000 \text{ fb}^{-1}$  assumed at 500 GeV (H-20 scenario)

- $\bullet~2{\times}1600\,fb^{-1}$  for LR and RL beam polarisation combinations
- $2 \times 400 \text{ fb}^{-1}$  for RR and LL beam polarisation combinations

assuming polarisation of  $\pm 80\%$  for electrons and  $\pm 30\%$  for positrons

arXiv:1903.01629

## CLIC

Total of  $5000 \, \text{fb}^{-1}$  assumed at  $3 \, \text{TeV}$ 

- 4000 fb<sup>-1</sup> for negative electron beam polarisation
- 1000 fb<sup>-1</sup> for positive electron beam polarisation

assuming polarisation of  $\pm 80\%$  for electrons



# Simplified DM model

Dark matter particles,  $X_i$ , couple to the SM particles via an mediator,  $Y_j$ .

Each simplified scenario is characterized by one dark matter candidate and one mediator from the set listed below:

	particle	mass	spin	charge	self-conjugate	type
DM	X <sub>R</sub>	$m_{X_R}$	0	0	yes	real scalar
	X <sub>C</sub>	m <sub>Xc</sub>	0	0	no	complex scalar
	$X_M$	$m_{X_M}$	$\frac{1}{2}$	0	yes	Majorana fermion
	X <sub>D</sub>	m <sub>X<sub>D</sub></sub>	$\frac{\overline{1}}{2}$	0	no	Dirac fermion
	$X_V$	$m_{X_V}$	1	0	yes	real vector
mediator	Y <sub>R</sub>	m <sub>Y<sub>R</sub></sub>	0	0	yes	real scalar
	$Y_V$	m <sub>Yc</sub>	1	0	yes	real vector
	T <sub>C</sub>	m <sub>Tc</sub>	0	1	no	charged scalar

# Backup slides



arXiv:2001.03011

# **Comparison with ILD study** Effective mass scale limits: $\Lambda^2 = \frac{M_Y^2}{|g_{eeY}g_{YYY}|}$

Limits from fast simulation (points) vs limits from full simulation (lines)



Very good agreement between full simulation and fast simulation results! ⇒ reliable extrapolation to low mediator mass domain...