$B \rightarrow \tau \nu_{\tau}$ and a hint for H+

Flavor Physics at the LHC ERA (workshop 2005-7,5 meetings, YellowReport in prep.)

Maria Krawczyk Collider and Astroparticle Physics Workshop 16.05.2007

Flavor Physics

Rare processes...

- Here inclusive and exclusive B to tau decays – data, SM and 2HDM
- W+ contra H+ :

in inclusive processes H+ exchange small correction, in exclusive – maybe 2 times bigger than W+!!! – in agreement with data

• Consider H+ in 2HDM (II) – large tan beta

Jeffrey Berryhill (FNAL) on behalf of other WG2 editors and contributors

- •Radiative penguin B decays ($b \rightarrow s, d \gamma$)
- •Electroweak penguin B decays (b \rightarrow s $l^+ l^-$)
- •Rare B decays with taus or neutrinos (B $\rightarrow \tau v$, b $\rightarrow s v v$)
- •Very rare leptonic B decays $(B_s \rightarrow \mu \mu)$



Photon is DIS probe of B (shape function, $m_{\rm B}$)

Broad sensitivity to new physics (2HDM, SUSY, LR, LED, little Higgs) Misiak et al.

$$\mathcal{B}(\bar{B} \to X_{s}\gamma)_{E_{\gamma}>1.6 \text{ GeV}} = \mathcal{B}(\bar{B} \to X_{c}e\bar{\nu})_{\exp} \left[\frac{\Gamma(b \to s\gamma)}{\Gamma(b \to ce\bar{\nu})}\right]_{\text{LO EW}} f\left(\frac{\alpha_{s}(M_{W})}{\alpha_{s}(m_{b})}\right) \times \\ \times \left\{1 + \mathcal{O}(\alpha_{s}) + \mathcal{O}(\alpha_{s}^{2}) + \mathcal{O}(\alpha_{em}) + \mathcal{O}\left(\frac{\Lambda^{2}}{m_{b}^{2}}\right) + \mathcal{O}\left(\frac{\Lambda^{2}}{m_{c}^{2}}\right) + \mathcal{O}\left(\frac{\Lambda}{m_{b}}\alpha_{s}\right)\right\} \\ \xrightarrow{\text{NLO}} \xrightarrow{\text{NNLO}} \sim 10\% \quad \sim 4\% \quad \sim 1\% \quad \sim 3\% \quad \sim 5\% \\ \xrightarrow{\text{perturbative corrections}} \text{non-perturbative corrections} \\ \xrightarrow{\text{non-perturbative corrections}} (\text{methods: Optical Theorem,} \\ \xrightarrow{\text{Operator Product Expansion,}} \\ \xrightarrow{\text{Heavy Quark Effective Theory}} \text{Need NNLO precision to compare with experiment}$$

Recent estimate of NNLO decay rate! **Inclusive** $b \rightarrow s \gamma$

In Misiakteh alith: H.M. Asatrian, K. Bieri, M. Czakon, A. Czarnecki, T. Ewerth, A. Ferroglia, P. Gambino, M. Gorbahn, C. Greub, U. Haisch, A. Hovhannisyan, T. Hurth, A. Mitov, V. Poghosyan, M. Ślusarczyk and M. Steinhauser.

hep-ph/0609232

 $\mathcal{B}(\bar{B} \to X_s \gamma)_{E_{\gamma} > 1.6 \text{ GeV}}^{\text{NNLO}} = (3.15 \pm 0.23) \times 10^{-4} \quad \textbf{7.3\% precision at NNLO}$

3% parametric $(\alpha_s(M_Z), \mathcal{B}_{\text{semileptonic}}^{\text{exp}}, m_c, \dots)$ **2.0% 1.6% 1.1%** $3\% m_c$ -interpolation ambiguity

3% higher order $\mathcal{O}(\alpha_s^3)$

P.S. :Becher & Neubert, hep-ph/0610067: -5% shift from 2-loop corrections at intermediate and soft scales $Br(\bar{B} \rightarrow X_s \gamma) = (2.98 \pm 0.26) \cdot 10^{-4}$

$\frac{\text{Inclusive b}}{\text{Experiment} - \text{NNLO Theory}} \rightarrow S \gamma$

B factories are likely to improve precision to 5%

= +1.2σ

Super B could push down E_{γ} cutoff from current 1.8 GeV to 1.5 GeV

Super B incl. ACP precision: 0.9 % @ 5 ab⁻¹ 0.3% @ 50ab⁻¹

Super B can measure incl. b \rightarrow d γ rate to 25% with 5 ab⁻¹





MSSM limits to be (re)evaluated: ideally NLO w/ minimal and general flavor violation

$B+ \rightarrow \tau^+ \nu$

Simple decay through weak annihilation



Sensitive to B decay constant f_B or to charged Higgs boson

$$\mathcal{B}(B_u \to \tau \nu)^{\rm SM} = \frac{G_F^2 m_B m_\tau^2}{8\pi} \left(1 - \frac{m_\tau^2}{m_B^2}\right)^2 f_B^2 |V_{ub}|^2 \tau_B = (1.59 \pm 0.40) \times 10^{-4}$$

 $tan^4\beta$ modifications in 2HDM II model:

$$R_{B\tau\nu} = \frac{\mathcal{B}(B_u \to \tau\nu)}{\mathcal{B}(B_u \to \tau\nu)^{\rm SM}} = r_H = \left[1 - \tan^2\beta \,\frac{m_B^2}{m_{H^\pm}^2}\right]^2$$

 $f_{\rm B}$ dependence can be removed via ratio with $\Delta m_{\rm d},$ error shrinks 25% \rightarrow <13% (Isidori & Paradisi)

$$\frac{\mathcal{B}(B_u \to \tau \nu)}{\tau_B \Delta M_{B_d}} \bigg|^{\rm SM} = 1.77 \times 10^{-4} \left(\frac{|V_{ub}/V_{td}|}{0.464}\right)^2 \left(\frac{0.836}{\hat{B}_{B_d}}\right)$$



$$B+ \rightarrow \tau^+ \nu$$

Belle result excludes (at tan β = 30) M(H+) < 100, 130 < M(H+) < 190 GeV



Probability Density Functions



MK'07

Current Constraints on 2HDM



MK'07

Current Constraints on 2HDM



Compare with inclusive decay

 $B \to X_c \tau \nu_{\tau}$



The decay rate is different from the exclusive case

$$R_{bc\tau\nu} = 1 - 2m_{\tau}^2 \frac{\tan^2\beta}{M_{H^{\pm}}^2} + m_b^2 m_{\tau}^2 \frac{\tan^4\beta}{4M_{H^{\pm}}^4}$$

Corrections small, only small x solution

$$x = \tan^2 / M_{H^{\pm}}^2$$

$$R_{B\tau\nu} = \frac{\mathcal{B}(B_u \to \tau\nu)}{\mathcal{B}(B_u \to \tau\nu)^{\rm SM}} = r_H = \left[1 - \tan^2\beta \,\frac{m_B^2}{m_{H^\pm}^2}\right]^2$$

Why?

- In the exclusive case –
- in SM helicity suppressed decay –
 so W- exchange prop. to m_tau/m_b
 in H+ exchange there is also prop to m_tau
 So, they are equally important and
 (1- X)² = 1 has two solutions X=0 and X=2
 - (X=x m_B^2)
- (H+ gives 2 times bigger than W contribution)

Conclusion

• SM is great, in full agreement with data

- Nevertheless new physics effects in rare may be large, even larger than the SM – tree level prediction
- Be careful